

# Proton Radiation Belt Remediation (PRBR)

Presentation to Review Committee

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# Removing Energetic Protons

- Removal is accomplished in the same way as HANE electron remediation: increase the pitch angle diffusion rate so that protons precipitate into the atmosphere
- Pitch angle diffusion rate is increased by producing waves with the proper wavelength to resonate with energetic protons
  - ULF waves in the 1 – 10 Hz band
- Unlike HANE electrons, inner belt protons are produced by very slow processes so remediation can be done periodically (e.g. for 1-2 years every 10 years) as well as monitored
- Remediation of natural inner belt protons would have an immediate operational impact
- **Environmental concerns**

# Proton Effects on Commercial Electronics

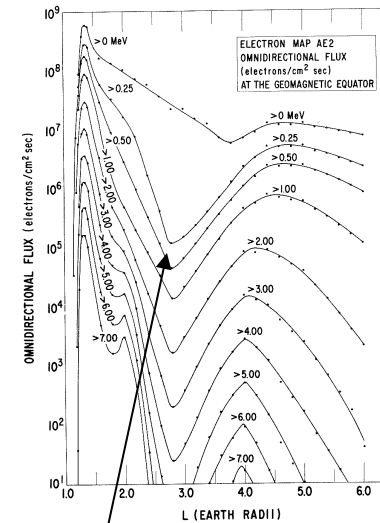
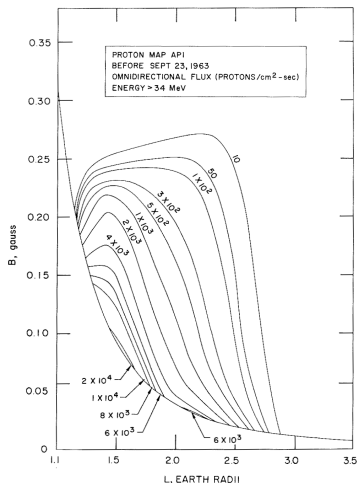
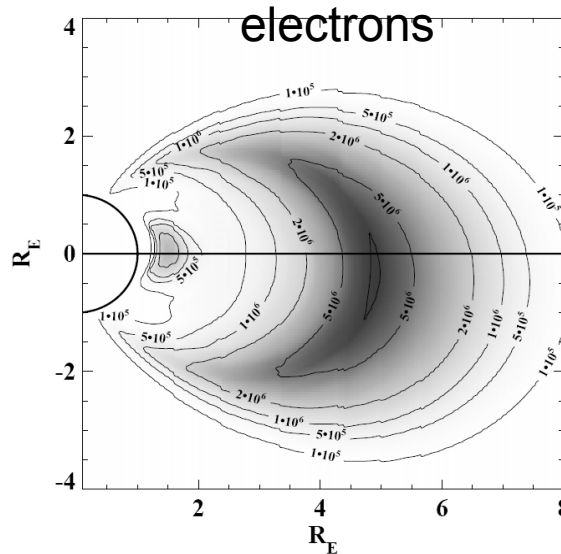
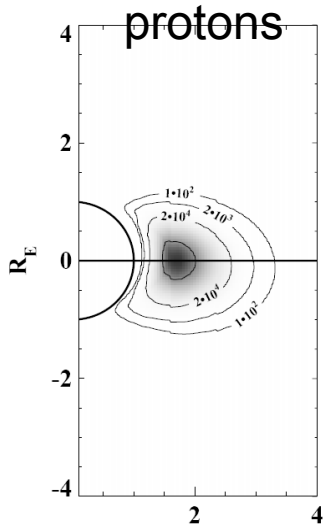
- Higher LEO Orbits: commercial electronics are regularly affected by proton upsets
- Lower LEO Orbits: affected during SAA transits and at high latitudes
- Example: IBM PowerPC 603 in Iridium (0.5 micron CMOS) – cache had to be disabled because of upsets caused by SAA

# Proton Effects Scaling with Feature Size

- As commercial feature sizes scale down, proton upsets will become much more frequent
- Critical charge for upset scales as (feature size)<sup>2</sup>
- For large feature sizes, protons cause upsets by hitting nuclei and releasing secondary particles that deposit charge
- At 65 nm and smaller, a proton deposits enough charge in silicon to cause an upset directly
- This can increase the proton SEU cross section by 2-3 orders of magnitude for deep submicron devices
- Major issue for micro-satellites

**IS PRBR A SOLUTION – ARE THERE ENVIRONMENTAL EFFECTS**

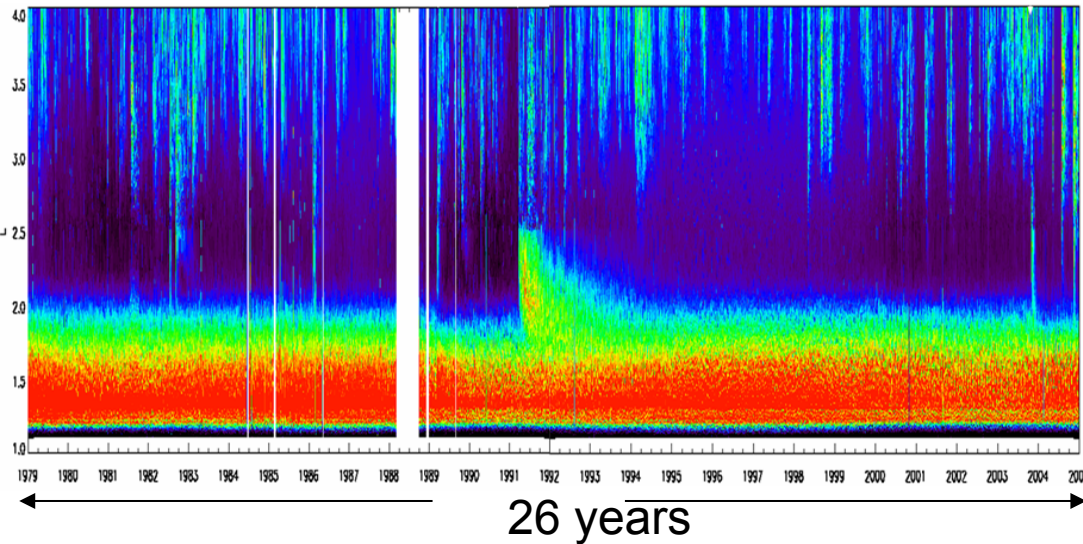
# Electrons (RBR) vs. Protons (PRBR)



- Only one proton belt – many MeV
- Peak near  $L \approx 1.8$
- Relatively stable
- **Sharp flux grad after  $L \approx 2$**
- **PRBR** focus on  $L=1.5-2.0$
- **Time scale of year(s)**

- Two electron belts with a slot region in between
- For  $E > 1$  MeV, peaks near  $L \approx 1.6$  and  $L \approx 4$
- **RBR** focuses on slot
- Time scale few days

# Proton Lifetime in the Inner RB



Steady State  $\rightarrow$  Source = Loss

Loss  $\rightarrow$  Slowing down by exciting and ionizing electrons in the thermosphere

$$T \approx 2 \times 10^4 (E / \text{MeV})^{1.3} (\# \text{cm}^3 / \langle \rho \rangle) \text{ years}$$

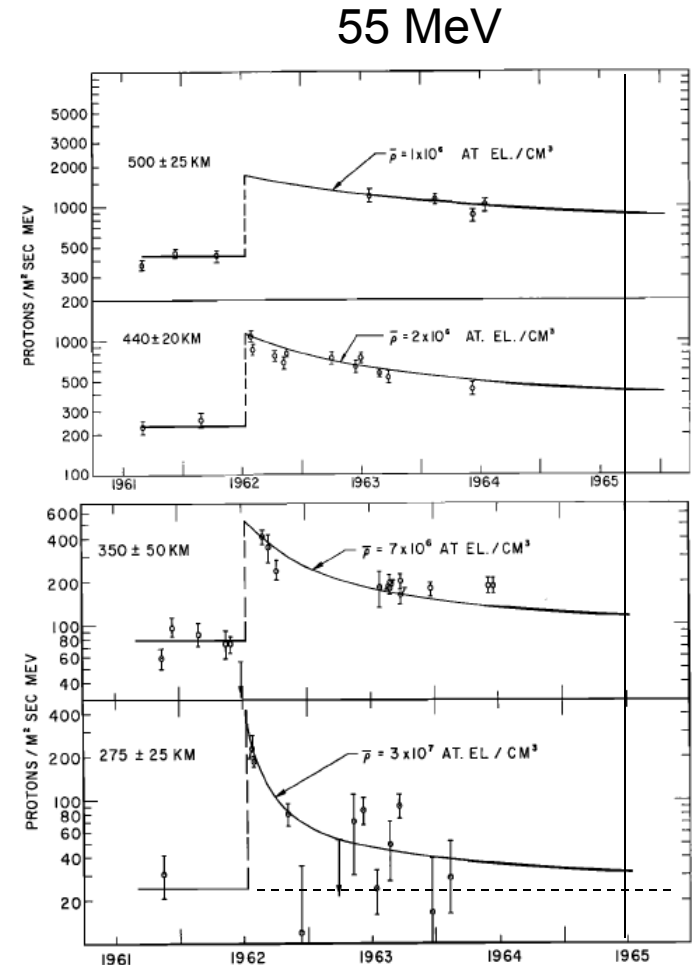


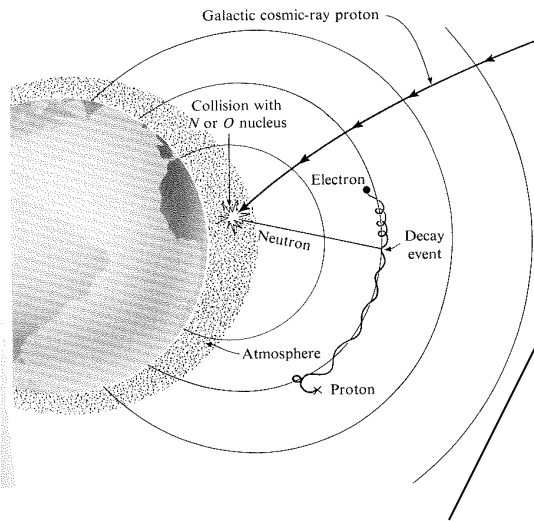
Fig. 10. Time variation of the 55-Mev proton flux for altitudes of  $H_s = 275, 350, 440,$  and  $500$  km from August 1961 to July 1964. The solid curves drawn for the period after July 1962 are the theoretical decay curves based on the first data point following July 1962.

# Proton RB Steady State Flux

$$\frac{dN(E)}{dt} = S(E) + \frac{d}{dE} \left( N(E) \frac{dE}{dt} \right)$$

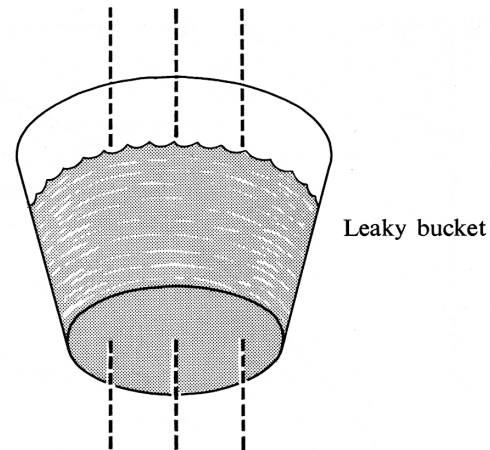
$$\frac{dE}{dt} = v \frac{dE}{dx} : \langle \rho \rangle$$

Slowing down by exciting and ionizing electrons of Oxygen atoms in the thermosphere

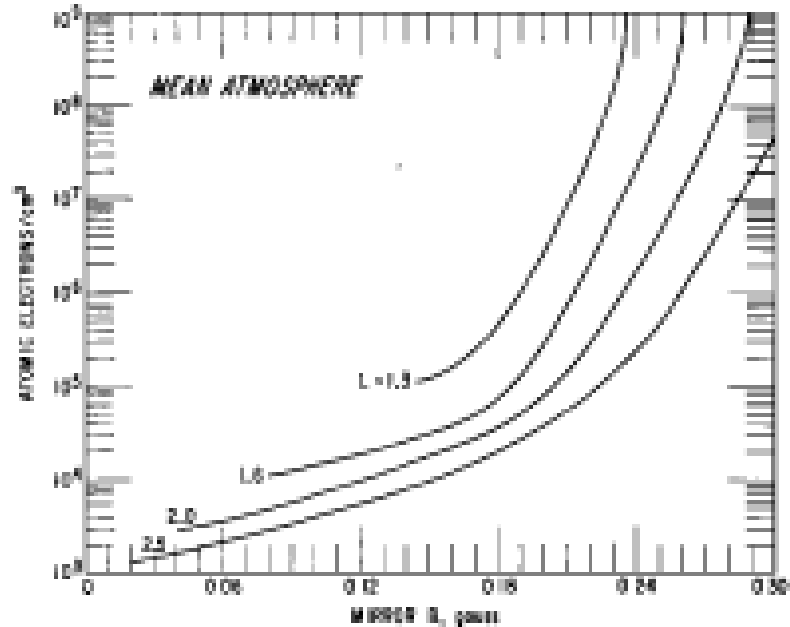
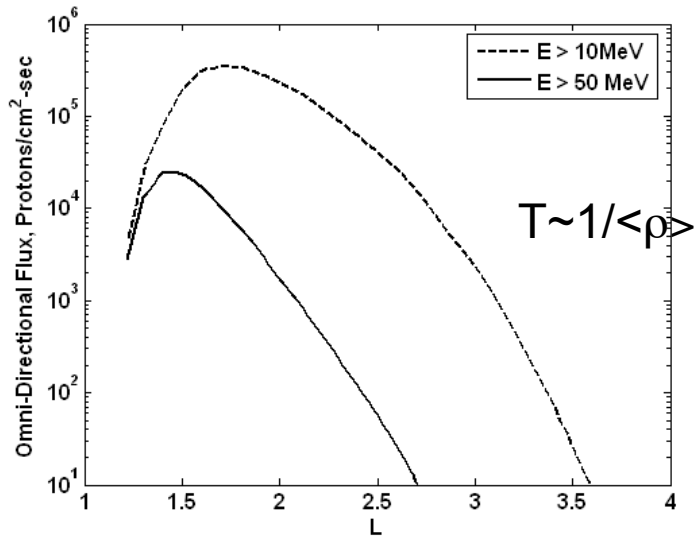


$$S(E) = - \frac{d}{dE} \left( N(E) \frac{dE}{dt} \right)$$

Cosmic Ray Albedo  
Neutron Decay (CRAND)



# Proton Lifetime vs. L



$$S(E) + \frac{d}{dE} \left( N(E) \frac{dE}{dt} \right) - \frac{N(E)}{T_L} = 0$$

$$\frac{d}{dE} \left( N(E) \frac{dE}{dt} \right) \ll \frac{N(E)}{T_L} \text{ if } L > 2$$

$$\frac{d}{dE} \left( N(E) \frac{dE}{dt} \right) \gg \frac{N(E)}{T_L} \text{ if } L < 2$$

If lifetime at L=1.5 is 20-50 years

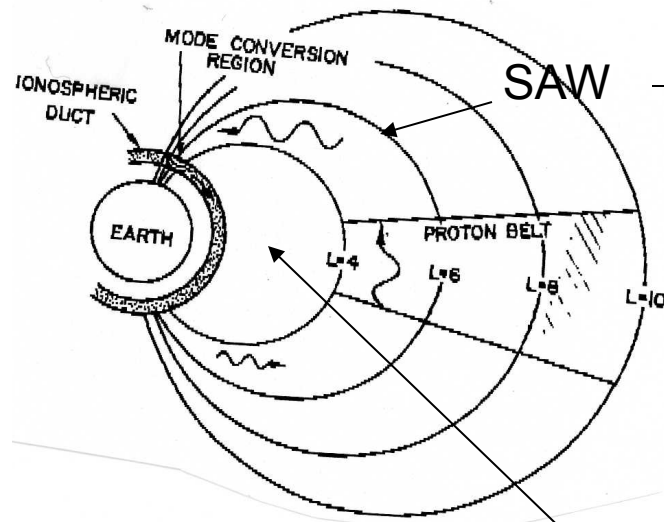
Lifetime at L=2.5-3 should be days

Need new loss mechanism

Pitch angle scattering into loss cone by SAW (A.Dragt) driven by Ion Cyclotron Instabilities



# Pitch Angle Diffusion (PAD) at high L

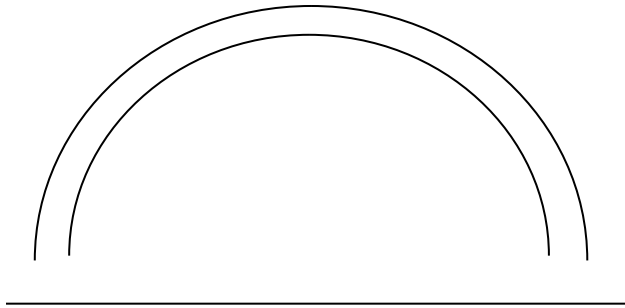


$$\omega - k_z v_z = \pm \Omega$$
$$k_z v_z \approx \Omega$$

Enhance proton loss rate in the inner RB by PAD on artificially generated and injected SAW

# Inner RB Proton Removal Requirements

Example for L=1.5  
Fill tube with SAW

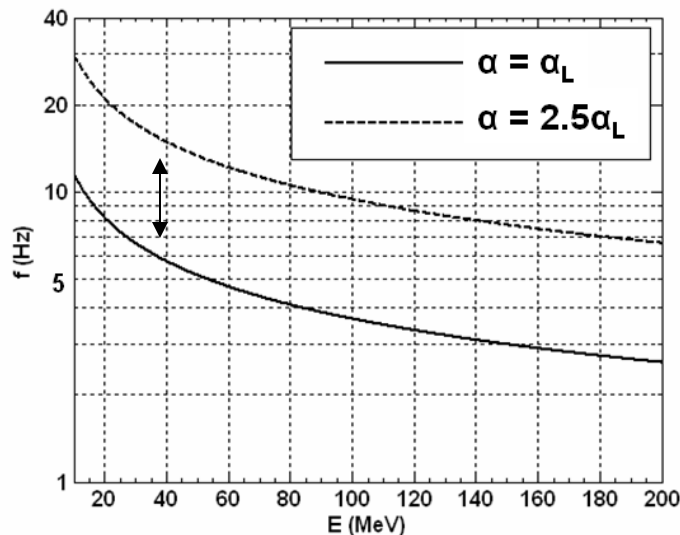


Frequency Selection for Resonance  
of Protons with SAW

$$\omega \approx k_z V_p$$

$$\omega = k_z V_A$$

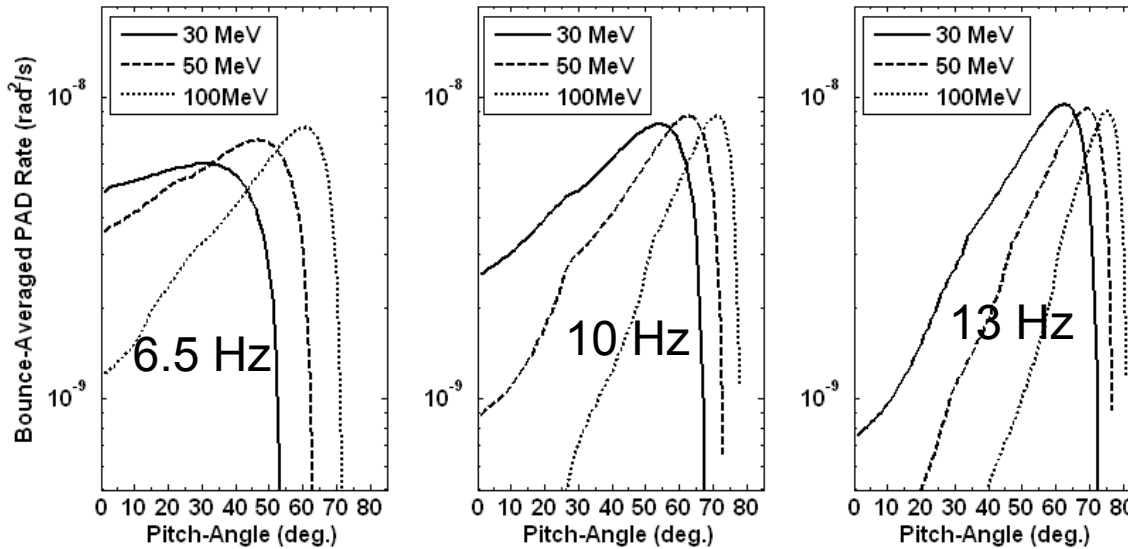
$$\omega(E, \alpha) \approx \frac{\Omega}{\cos \alpha} \sqrt{\frac{M V_A^2}{2E}}$$



Frequency requirement for equatorial  
resonance with SAW at L=1.5

Frequency range 5-15 Hz

# PAD Rate and Lifetime



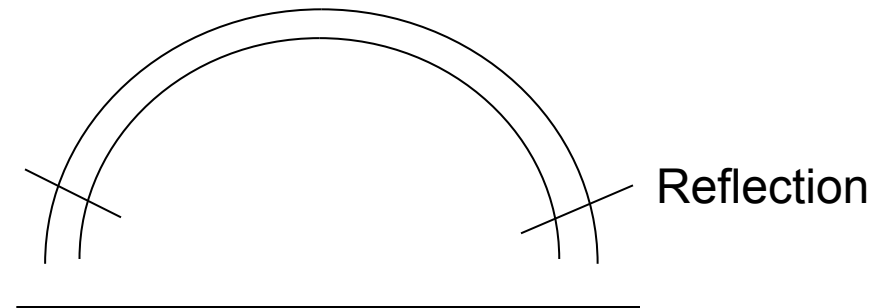
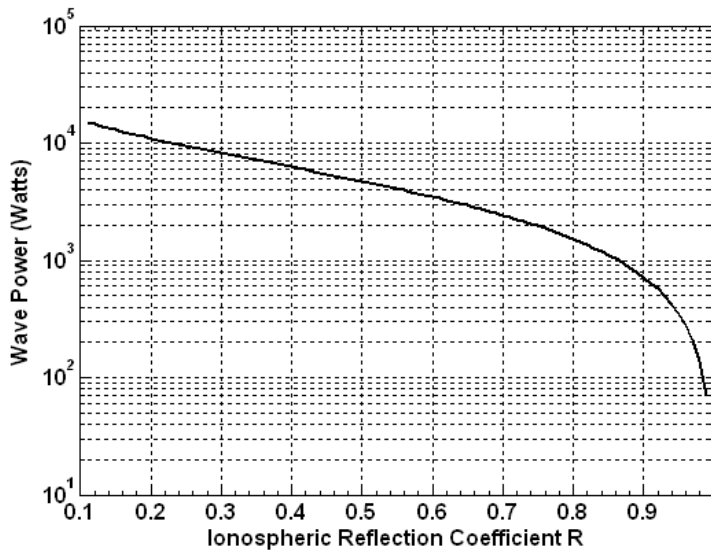
$$\Delta f/f \ll 1/2, \langle B \rangle \ll 25 \text{ pT}$$

	$f_1 = 6.5 \text{ Hz}$	$f_2 = 10 \text{ Hz}$	$f_3 = 13 \text{ Hz}$
E = 30 MeV	1688 days	880 days	595 days
E = 50 MeV	900 days	586 days	920 days
E = 100 MeV	580 days	1032 days	1600 days

Energy stored in SAW at  
 $L=1.5$  and  $\Delta L=.1$  with  $\langle B \rangle$   
 $=25 \text{ pt is}$

$$W = 75 \text{ kJ}$$

# Injection Power Requirement



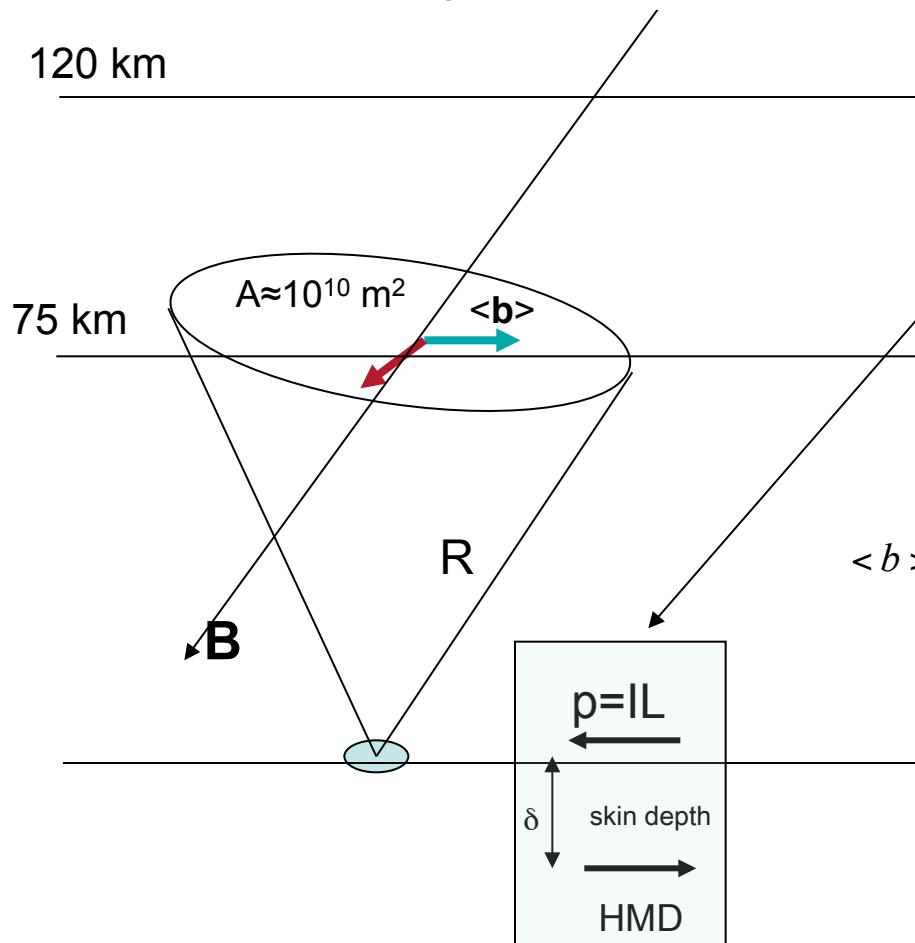
Injection power required to maintain  
75 kJ at  $L=1.5$  per .1 L width

# How to Get these Waves - Ground-based Transmitter Options

- Initial estimates indicate that less than kWatt level of ULF injected into the  $L = 1.5-1.8$  region is required to get interesting removal lifetime
- What does it take to get it
- There are a number of potential options:
  - Conventional ULF/ELF transmitters (grounded dipoles)
  - Rotating electromagnets (conventional and low and high temperature superconducting)
  - Space based rotating magnets or neutral gas injection
  - Electrojet-free F-region ionospheric heating
- Present estimates for the first two

# Ground-based Conventional Transmitter - HED

To inject 1 kW we require  $\langle b \rangle \approx 30$  pT at 75 km, the bottom of the magnetized ionosphere.

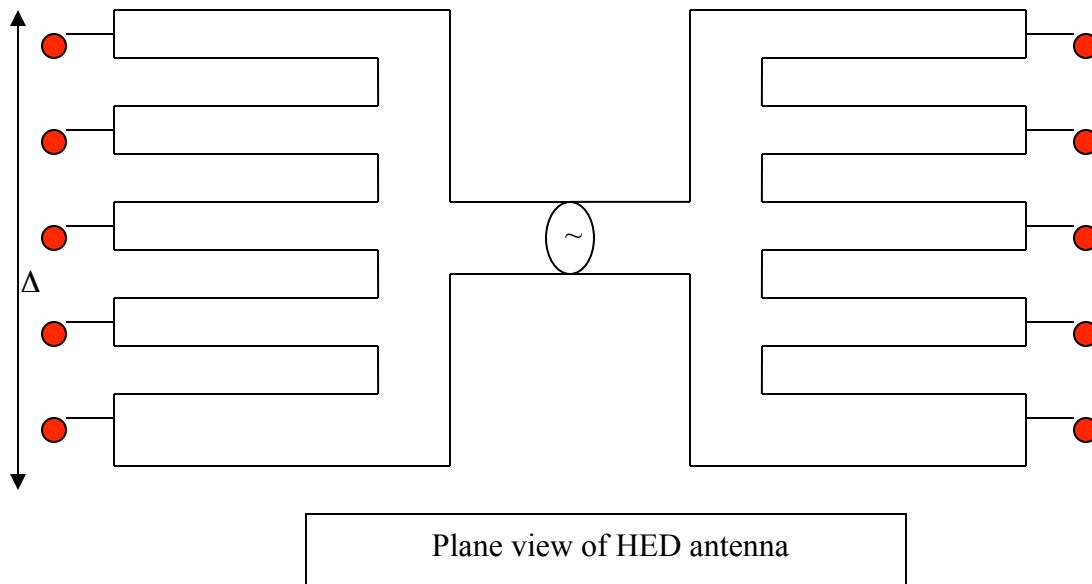


Conventional ULF/ELF sources (like Sanguine/Seafarer) are grounded wires – HED (Jason Study by Perkins et al.)

$$\langle b \rangle \approx 60(IL^2 / 2 \times 10^{10}) \left[ \frac{\delta}{2(L + \delta)} \right] pT \approx 30(IL^2 / 2 \times 10^{10}) pT$$

Need  $M \approx IL^2 \approx 2 \times 10^{10} \text{ A-m}^2$

# Traditional HED Sources



$$M_{eff} = \frac{Il^2\delta}{2(l+\delta)}$$

$$B \approx \mu \frac{M_{eff}}{h^3} \approx \frac{\mu Il\delta}{2h^3} \frac{1}{(1+\delta/l)}$$

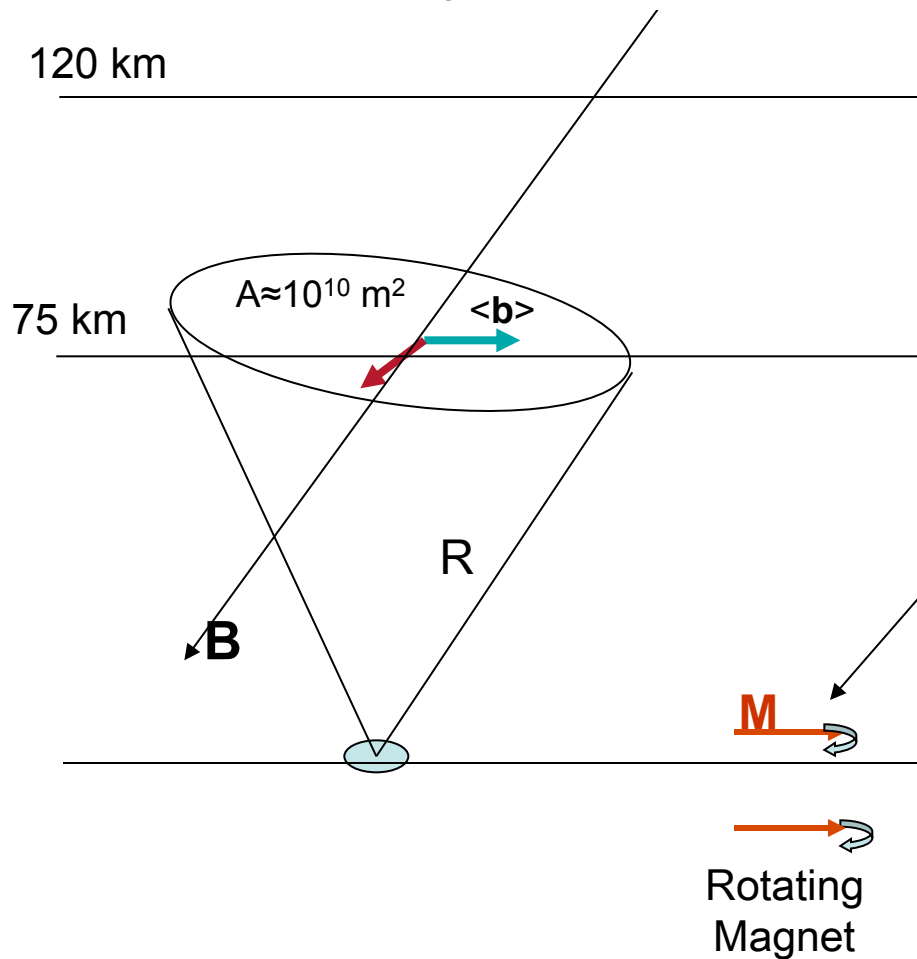
$$I = 3.3(P/MW)^{1/2} (\sigma/10^{-2})^{1/2} (L/km)^{1/2} \text{ kA}$$

$$Il^2 = 2 \times 10^{10} (P/MW)^{1/2} (\sigma/10^{-2})^{1/2} (L/2km)^{5/2} A-m^2$$

Jason study Perkins et al.

# Ground-based Rotating Magnet Array (RMA)

To inject 1 kW we require  $\langle b \rangle \approx 30$  pT at 75 km, the bottom of the magnetized ionosphere.



A superconducting magnet rotating at a ULF frequency has a reflected image in phase with primary

$$\langle b \rangle \approx 30(M / 5 \times 10^9) pT$$

Need RMA with  $M \approx 5 \times 10^9$  A-m<sup>2</sup>



# Innovative Sources: Rotating Magnets

- Rotating superconducting magnets are useful for frequencies of up to 10 Hz
- They are compact sources of large moments and can be used in arrays
- Example design:
  - Superconducting coil 5 m high x 5 m wide x 5 m long
  - 25 m<sup>2</sup> area
  - 100 Amps DC current
  - $4 \times 10^4$  turns
  - $M = 10^8$  A-m<sup>2</sup> per coil, meaning 50 coils are needed
- Cost estimate: ~\$1M/coil
  - LTS wire at \$2/kA-m: \$160k/coil
  - Dewar and refrigeration: \$500k/coil (assuming LHe large plant shared across dozens of coils)
  - Mechanical rotation: \$300k/coil (depends strongly on maximum frequency)

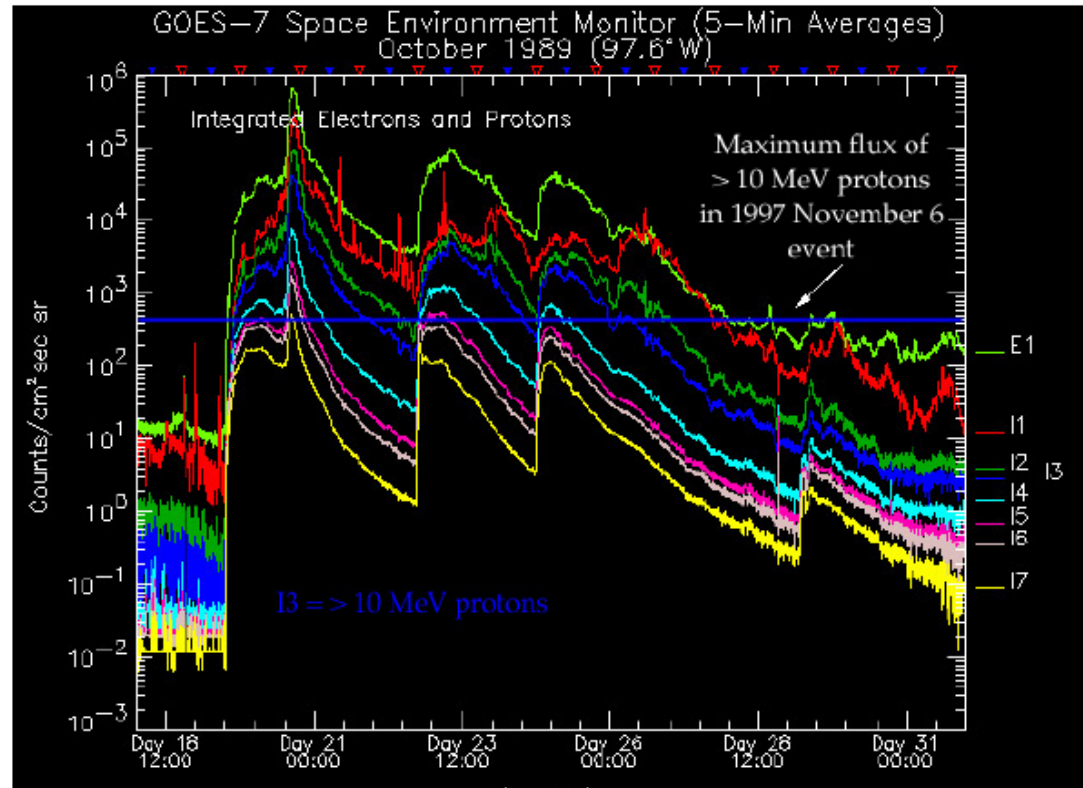
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- Unlike HANE electrons, inner belt protons are produced by very slow processes so remediation can be done periodically (e.g. for 1-2 years every 10 years) as well as monitored
- Remediation of natural inner belt protons would have an immediate operational impact
- Similar ULF system could potentially be used for MeV electrons

# Environmental Effects of Energetic Proton Precipitation in Middle Atmosphere

## Solar Proton (SPE) Events associated with CMEs

Flux of  $> 10$  MeV protons  
 $> 10^4$  #/cm<sup>2</sup> sec  
Leads to 20% variation of Ozone in middle atmosphere (40-50 km) with recovery time of week



Verronen et al JGR, 2005

Jackman et al. JGR, 1995

Noise for  $E > 10$  MeV 10 #/cm<sup>2</sup>sec

# Environmental Issues – Effects in the Stratosphere – Ozone Loss

PRBR precipitation flux

$L=1.5-2$

Precipitation area  $10^{17}$   
 $\text{cm}^2$

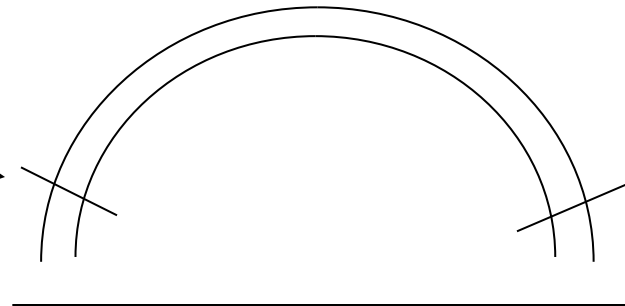
Proton inventory  $10^{23}-10^{24}$

Assume all particles  
precipitate over one year

Proton flux in atmosphere

$<1 \text{ \#/cm}^2 \text{ sec}$

Smaller than noise



# Environmental Effects - Magnetic

- Diamagnetic current due to Inner Belt protons  $\ll$  Ring Current
- Magnetic Moment of Ring Current  $\ll$  Magnetic Moment of Earth ( $7 \times 10^{21} \text{A}\cdot\text{m}^2$ )
- Magnetic effect from PRBR negligible